Final State Interactions (GENIE) S. Dytman Univ. of Pittsburgh

- 1. Typical nuclear model in event generators
- 2. Hadron (pion, nucleon) fsi



Simulation

- Event generators (e.g. GENIE) do full vA interaction
- Basic nuclear model is Fermi Gas
 - ✓ Single particle densities right, NN correlations wrong (except SRC)
 - ✓ Medium effects approximated (Fermi momentum, binding genergy)
- Basic interaction model is *intranuclear cascade (INC)*
 - ✓ All particles are free (corrected)
 - \checkmark Many final states can be described



vp or vn \rightarrow form. zone \rightarrow fsi



Fermi gas model nucleons are INDEPENDENT

- Justified by (e,e') data of ~1970.
- Smith, Moniz (1972)
- Bodek, Ritchie (1981)
 - Assume struck nucleon off-shell, outgoing nucleon on-shell (fsi issue!)
- Basis for all ν event gen.
 - ✓ Fermi mom. good,
 - ✓ Pauli blocking good, but....
 - ✓ How to handle binding energy?



μ



nuclear structure

- In *Fermi Gas* model, nucleons don't interact.
- They are bound in a potential, momentum and energy disconnected, E²≠p²+m² (off-shell).
- *Structure function* (Benhar) uses potential to calculate probability for qe as a function of mom and energy
- Interactions produce *correlations* which effect data.
- Can be used for electrons or neutrinos, but each nucleus is different!
- This is only qe, application Feb $^{6}_{to}$ $^{2009}_{1\pi}$ production impossible.





v CC interaction Diagrams

`normal' qe

$$\sum_{n < F} \bigvee_{W^+}^{p} n^2$$

'normal' π prod (+fsi)



qe with NN correlations Called MEC, SRC (all Q²)



RPA correlations at $Q^2 < 0.2 GeV^2$.



5



Application to (e,e') I

Calc of Benhar Uses structure function Note that data is incl xs vs. ΔE at fixed E_0 , θ . Calc of O'Connell, Sealock. Shows separate effects of true qe, correlations, π production.







Application to v data (Tina Leitner)





v-nucleus calculations

- (e,e'), CVC \rightarrow Vector (v, μ), get Axial from PCAC
- FG has larger cross section, more peaky
- No data for comparison!









Historical perspective

• Electron scattering

 \checkmark eN cross sections well-known early

- ✓ Dipole approximation important organizing principle
- \checkmark eA data used to learn about nuclear structure
- Neutrino scattering
 - ✓ vN cross sections moderately well-known, use calc of Llewlyn-Smith
 - \checkmark Dipole form factor important (M_A)
 - ✓ Must use vA data to measure M_A , must assume knowledge of nuclear structure and reaction mech.



Final state interactions (FSI)

- Historically most difficult part of any nuclear simulation code
- 2nd significant change between nuclei (nuclear structure!)
- INC model is `simple', able to describe many final states important to vA interactions.
 - ✓ $\nu C \rightarrow \mu^- p \text{ vs. } \nu C \rightarrow \mu^- ppn \text{ vs. } \nu C \rightarrow \mu^- ppppnnnnn$
 - ✓ Describe NC coherent π^0 production in nuclear medium
 - ✓ Describe CC processes in nuclear medium, e.g. pi production followed by absorption (important background).
- Interaction probability by mean free path (mfp)
 - ✓ $\lambda(r,E)=1/[\sigma_{\pi N}(E)*\rho(r)]$
 - Use charge density from (e,e), πN and NN xs from GWU
 <u>http://gwdac.phys.gwu.edu/</u>

Feb 6, 2009 Fob(interaction)=1-exp(-x/ λ)



- Ideally, we'd have lots of vA data with all final state particles identified. We have 1 bubble chamber experiment with ~1000 events (Merenyi). (neutral pions, neutrons hard)
- We do have lots of π[±]A, pA, nA, and γA data which measure the same properties. Use them until SCIBoone, Minerva... data available. *Simulation is key now!*



overview

- Hadronic final state interactions (fsi) matter
- v interacts through weak interaction (λ~ly), but
 p, n (N) emitted, π produced (strong interaction, λ~Fm)
- Therefore, ~10-30% of particles in final state come from fsi, not the primary interaction!
- PROBLEM: fsi can mask the primary interaction, e.g. π production followed by π absorption appears as qe event!





Results from Jim Dobson

- GENIE simulations
- Top plot is v_{μ} Fe, 1 GeV
- Table for ν_{μ} O, T2K beam.



Final-				Prir	nary Hae	dronic Sy	stem			
State	$0\pi X$	$1\pi^0 X$	$1\pi^+X$	$1\pi^-X$	$2\pi^0 X$	$2\pi^+X$	$2\pi^-X$	$\pi^0\pi^+X$	$\pi^0\pi^-X$	$\pi^+\pi^-X$
$0\pi X$	293446	12710	22033	3038	113	51	5	350	57	193
$1\pi^0 X$	1744	44643	3836	491	1002	25	1	1622	307	59
$1\pi^+X$	2590	1065	82459	23	14	660	0	1746	5	997
$1\pi^-X$	298	1127	1	12090	16	0	46	34	318	1001
$2\pi^0 X$	0	0	0	0	2761	2	0	260	40	7
$2\pi^+X$	57	5	411	0	1	1999	0	136	0	12
$2\pi^- X$	0	0	0	1	0	0	134	0	31	0
$\pi^0 \pi^+ X$	412	869	1128	232	109	106	0	9837	15	183
$\pi^0\pi^-X$	0	0	1	0	73	0	8	5	1808	154
$\pi^+\pi^-X$	799	7	10	65	0	0	0	139	20	5643



Alternatives (fsi)

• Quantum mechanical model

- ✓ Hadron wave effects, correlations done correctly
- ✓ We know QM essential for proper treatment of nucleus
- \checkmark Impossible to calculate multiple particle final states properly
- \checkmark Propagating hadrons tend to remain on-shell (not π abs)

• GIBUU (semi-classical model)

- \checkmark Giessen group reinvigorated interest
- ✓ Many applications, Tina Leitner will present it next week.
- \checkmark Computing needs intensive compared to INC

• Limits of INC should be understood!

- \checkmark Comparison with hA data excellent start
- \checkmark Comparison with (e,e') data is essential for nuclear structure.



Applicability of INC

To ensure h sees only 1 nucleon at a time, we want $\Lambda <<\lambda << R$ and $d <<\lambda$.

 Λ = pion size λ = pion mean free path R= nuclear size d= nucleon spacing

DICEY!!! (but it works!!)





What is INC formalism good for?

- Inelastic reactions, esp. particle production processes.
- Only pion induced reactions shown here, but still some impressive examples.





Mashnik INC calcs

State of the art code, under development for 'decades'.
π inclusive cross sections at T_π~500 MeV show many effects. ✓Quasielastic scattering ✓Pion production
Here, they examine the effect of a 2Fm/s hadronization time.







CEM03.01 vs. p ⁵⁶Fe data (*their tests*)

20

15

10

5

0

10²

Wide range of final nuclei likely!



One final state (⁵²Mn) as function of proton energy Data has systematic error troubles Many calculations shown, CEM03.01 does best overall (prediction).

10³

Proton energy (MeV)



What is downside?

- No quantum mechanics [but no qm model usable]
- Nuclear structure is Fermi gas.
- Unlikely to do well for elastic processes which are typically diffractive in nature.
- Papers refer to model choices which may be covering up problems (sometimes hard to tell).



- Shapes are very similar ($\rho \sim 0.16 \text{ N/Fm}^3$)
 - ✓ nuclei density saturated (same as neutron star)
 - ✓ Woods-saxon distribution describes all A≥20

(R~1.4 x A^{1/3}), e.g. 1.4*3.8=5.3 Fm for Fe

✓ Modified Gaussian describes all A<20</p>







Nuclear density (GENIE can do almost all nuclei)

- Use data values for common nuclei, interpolate for others
 - $\checkmark\,$ Gaussian for ⁴He, modified Gaussian for ¹²C, ¹⁴N, and ¹⁶O
 - ✓ Interpolate to others for A≤20
 - \checkmark 2 param Woods-Saxon for A>20, data for ²⁷Al, ²⁸Si, ⁴⁰Ar, ⁵⁶Fe, ²⁰⁸Pb
 - ✓ Interpolate for others (errors are few %)
- We empirically add to nuclear size
 - $\checkmark~0.5^*\lambda_{deB}$ Fm for nucleons, $1.0^*\lambda_{deB}$ Fm for pions (v 2.4.0)
 - $\checkmark\,$ Empirically, this gets good agreement with vA, πA and pA data
 - Theoretically, this is justified because hadrons have size ~ 1 Fm





Nuclear systematics





Reaction glossary

- Elastic $\pi^{+12}C \rightarrow \pi^{+12}C 1\pi^{+11}$ in final state (E>0.8E₀)
- cex $-\pi^{+12}C \rightarrow \pi^{012}N 1\pi^{0}$ in final state (E>0.8E₀)
- Inelastic $\pi^{+12}C \rightarrow \pi^{+12}C 1\pi^{+11}$ in final state (E<0.8E₀)
 - $\pi^{+12}C \rightarrow \pi^{+11}B p 1\pi^{+11} n final state (E<0.8E_0)$
- Absorption $\pi^{+12}C \rightarrow nnp^{10}C 0\pi^{+}$ in final state
- Total = sum of all
- Reaction = sum of all except elastic
 - \checkmark can be well-described by INC
- Elastic scattering is wave property (not in INC)
- Quasifree = reaction with nucleus looks just like 2-body +Fermi motion



Quasifree (QF) reaction mechanism important...

•QF means hadron interacts with nucleons in nucleus as though they were free (with momentum) •INC calcs by Fraenkel (1982) agree with data!







Role of QF mechanism in πA inclusive scattering



...but far from complete (FSI!)



	Derra Deta	$30{ m MeV}$	Extrapolated		
	Kaw Data	Threshold	to $0\mathrm{MeV}$		
5p	0.013 ± 0.001	0.04 ± 0.01	0.64 ± 0.13		
4p	1.11 ± 0.10	2.0 ± 0.2	5.1 ± 1.0		
$_{3p}$	19.9 ± 1.2	26.8 ± 2.5	28.4 ± 4.0		
$_{\rm 3pn}$	2.0 ± 0.2	11.9 ± 1.3	33.2 ± 7.5		
2p	69.8 ± 4.2	72.9 ± 5.8	$43.6 \pm 5.2 \leftarrow$		
$_{2p1n}$	11.9 ± 0.9	62.9 ± 6.6	$75. \pm 10.$		
2p2n	0.67 ± 0.05	5.6 ± 1.0	$21. \pm 8.$		
2 pd	9.2 ± 1.0	10.3 ± 1.2	7.9 ± 1.4		
pd	14.6 ± 2.3	9.8 ± 1.7	4.2 ± 1.0		
pdn	3.0 ± 0.4	13.8 ± 2.4	10.6 ± 2.5		





Components of INC

- Decide where an interaction occurs
 ✓ Mean free path λ(E,r)=1/[ρ(r)σ(hN)] makes sense
 - \checkmark Choose int point from exponential distribution
- Decide which interaction to simulate
 - ✓ hA and hN differ here✓ Both tied strongly to data





mean free path in iron (solid=old, dashed=new



GENIE fsi models

- Goal describe hadron-nucleus reactions $T_h < 1$ GeV, all nuclei
- hA (today, discussed by Costas Tuesday)
 - \checkmark Introduced in 2006, improvements in 2008
 - \checkmark Schematic, but tied strongly to data $~\pi$
 - \checkmark Originally intended for systematic errors
- hN (tomorrow)
 - \checkmark About to be introduced
 - ✓ Full INC code for π^+ , π^- , π^0 , p, and n
 - \checkmark Surprisingly fast and accurate
- Results of 2 codes suggest advances for each other





GENIE models - hA

- hA is only FSI model in GENIE 2.4.0
 - ✓ At most 1 FSI
 - ✓ If FSI, choose final state according to total cross section (e.g. π absorption total xs is ~25% of total)
 - \checkmark Use data for iron, extrapolate ${\sim}A^{2/3}$ for all others
 - ✓ Use pre-existing code for angular distributions (wrong)
 - ✓ Error bars in plots used for systematic error studies $_{\pi^{+56}Fe}$



Mean free path





p ⁵⁶Fe scattering, implementation

Use total xs of CEM03, focus on aspects important for MINOS
Total reaction cross section governs total amount of stuff coming out
Lack of data still an issue, esp. at KE>800 MeV.

•Use phase space for angle, energy distributions, unlikely to be large



•Calculations built up over many years, reliability appears to be very high.

•Code is huge, deals with phenomena in which we have no interest (10n8p knockout)

•Implementation can cause errors

No n calculations (differences small)

Some π production channels incomplete

INC can't do elas scat well, need separate code (future).
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INTRANUKE hA strategy

- Use CEM03.01 calculations for p ⁵⁶Fe except optical model for σ_{elas} (scale by A^{2/3} for other nuclei)
 Results at 50, 100...1000 MeV incident energy
- •Channels included:



Elastic	p ⁵⁶ Fe→p ⁵⁶ Fe
Charge exchange	p ⁵⁶ Fe→n ⁵⁶ Co
Inelastic	p ⁵⁶ Fe→p´ ⁵⁶ Fe
Breakup	p ⁵⁶ Fe→pn ⁵⁵ Fe
	p ⁵⁶ Fe→pp ⁵⁵ Mn
	p ⁵⁶ Fe→ppn ⁵⁴ Mn
	p ⁵⁶ Fe→pnn ⁵⁴ Fe
Breakup-generic	p ⁵⁶ Fe→pppnn ⁵² Cr
Pion production	p ⁵⁶ Fe→π ⁺ n ⁵⁶ Fe
	p ⁵⁶ Fe→π⁺ π ⁰ n ⁵⁶ Fe



π rescattering tougher

- 3 charge states, no π^0 data
- Data for total cross sections up to KE~1 GeV
 - \checkmark Data for component final states largely at KE<400 MeV
 - ✓ Angle, energy distribution data not extensive (like pA)
- CEM03 has real problems (~40% discrepancies)
 ✓ Use only for guidance at high KE
- Use data (~10-20% errors) at low energy, CEM03+intuition at high energy.
 - ✓ Recently discovered inclusive p, n prod data will be very useful.
 - ✓ Elastic xs must decrease rapidly with energy
 - ✓ Inelastic xs must rise
 - ✓ Total remains largely constant



TOTAL cross section taken from data

✓ Quality data for many targets (Carroll, et al.) at T_{π} <450 MeV ✓ Quality data for light targets (Clough, et al.) at T_{π} <860 MeV ✓ Note ~flat energy dependence for all targets at high energy ✓ Use power fit to A dependence to extrapolate



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CEM03.01 - π^{+56} Fe [σ_{RFAC}]

•REACTION (abs+inel+cex=tot-elas) cross section trickier ✓ Poor agreement vs. Ashery et. al data at low T_{π} . ✓ Expect σ_{elas} small high T_{π} . \checkmark Compensating factor has too much energy dependence. \checkmark Use data with \sim constant extrapolation (like calc. and

total xs)





CEM03.01 - π + 56Fe [σ_{ABS}]

- •Total ABSORPTION cross section is tricky (like reac)
- •Ashery et al. data at low energy, nothing at higher energy
- •Use calculation as a guide to xs at high energy
- •Use data + extrapolation for total absorption xs
- •Use final states from CEM03 in same ratio.
- •Ashery data has est. error ~20%





CEM03.01 – π + 56Fe [σ_{π^0}]

- •Total Inclusive π^0 cross section even worse [fortunately small]
- •2 old data points (1 from LAMPF), nothing at higher energy
- •CEM03 calc. rises at high energy (*pion production*)
- •CEM03 calc. divided by 2 agrees with data at low energy
- $\sigma(\pi^- p \rightarrow \pi^0 n)$ peaks at res, $\sigma(\pi^- A \rightarrow \pi^0 A)$ should be prop. to this
- •Use scaled $\sigma(\pi^-p \rightarrow \pi^0 n)$ (purple) for $\sigma(\pi A \rightarrow \pi^0 A)$ at all energies
- •Use CEM03/2- $\sigma(\pi^+A \rightarrow \pi^0A)$ for $\sigma(\pi^+A \rightarrow \pi^0\pi^+A)$ at T_{π}>300





CEM03.01 - π + 56Fe [$\sigma_{inel}, \sigma_{elos}$]

•We now have 2 checks,

- $\sigma_{\text{elas}} = \sigma_{\text{tot}} \sigma_{\text{reac}} (\sigma_{\text{reac}} = \sigma_{\text{abs}} + \sigma_{\text{inel}} + \sigma_{\text{cex}})$
- $\sigma_{\text{inel}} = \sigma_{\text{reac}} \sigma_{\text{abs}} \sigma_{\text{cex}}$

•Compare results with Ashery data at low energy

✓ Have to trust CEM03 at high energy

After a little playing, it works

Everything is consistent,

though not unique.





hA strategy $[\pi]$

- •Mix of data, intuition, CEM03.01 calc. for π ⁵⁶Fe (scale by A^{2/3} for other nuclei)
- •Jumps of σ_{abs} at low energy is in data!
- •More adjustment needed, but basic strategy is done



Elastic	π^{+56} Fe $\rightarrow \pi^{+56}$ Fe
Charge exchange	π^{+56} Fe $\rightarrow \pi^{056}$ Cr
Inelastic	π^{+56} Fe $\rightarrow \pi^{+1}$ N ⁵⁶ Fe
Absorption	π ^{+ 56} Fe→ pn ⁵⁴ Fe
	π^{+} ⁵⁶ Fe $ ightarrow$ pp ⁵⁴ Mn
	π^{+} ⁵⁶ Fe \rightarrow ppn ⁵³ Mn
	π^{+} ⁵⁶ Fe $ ightarrow$ pnn ⁵³ Fe
Abs-generic	π^{+} ⁵⁶ Fe \rightarrow ppnn ⁵² Mn
Pion production	π^{+56} Fe $\rightarrow \pi^{+}\pi^{0.56}$



Distributions of fsi particles

- Scattered particles
 ✓ Isotropic [wrong]
- Particles produced (e.g. π absorption)
 ✓ Phase space
- Does it matter?



Old NEUGEN vs. other models

Results of study of H. Gallagher and others for Nuint04(ν_μp)
ν_μp → μ⁺pπ⁻ was main source of protons in Neugen, no baryon rescattering, no pion absorption (added in 2005)





Compare to 2005, old (5 GeV neutrinos)





Compare hA to others



LEN 0, 7003



 π^+ at 5 GeV

Not much changes despite new mechanismsOther models, 2005, and hA in agreement





Strategy for est. errors

•Use error bar of data (KE<400MeV)
•At higher energies,
>Error bar*1.5 for σ_{tot}
>Error bar*2 for others





Compare to 2005, old (5 GeV neutrinos)







רכט ט, ∠טטש



hN Validation process

 Test mean free path with total cross section
 Test reaction processes with component total cross sections and inclusive distributions.





Inclusive distributions

No previous effort made to match complete theory or these data





Caveats, future

- Problems with hA (all fixed with hN)
 - ✓ π^+ and π^- are identical (in fact, π^- interacts a little more)
 - \checkmark Only works with N=Z nuclei (Pb will be somewhat wrong)
 - \checkmark Angular distributions are wrong
- New hN model
 - \checkmark Full INC calculation for pions, nucleons
 - ✓ Build hA interaction from hN data (phase shift data)
 - \checkmark Extensive testing almost complete, will be in v2.6.0 (soon)
 - \checkmark Can then do some fixes in hA